

The challenging application of cosmogenic dating methods in residual glacial landforms: the case of Sierra Nevada (Spain)

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Abstract

An accurate review of the literature on surface exposure dating methods shows evidence of the difficulty in applying cosmogenic dating methods to old moraines because of the intensity of late Quaternary erosion processes. Moreover, as in some previous cases, we also found special difficulties in applying these methods to LIA moraines, caused by the intensity of current paraglacial processes. The objective of this study is to apply cosmogenic dating methods to very old and very young moraines, which in both cases have been or are being affected intensively by erosion. With this purpose, we collected samples of boulders from moraines corresponding to (i) the penultimate glaciation and (ii) the Little Ice Age (LIA), both from Sierra Nevada in the south of the Iberian Peninsula. The sampling strategy was based on a preliminary accurate analysis of the geomorphological settings of two valley sites that resulted in the collection of only four boulder samples from an old moraine and three more from a very recent moraine. Using in situ-produced cosmogenic ¹⁰Be to date these boulders, the old samples yielded an age of ca. 130–135 ka for moraine stabilization. The younger samples indicate that the LIA moraine accretion probably occurred between the fourteenth and seventeenth centuries, with a subsequent stage of accumulation during the nineteenth century as suggested by historical documents. Dating a glaciation that occurred prior to the last Pleistocene glacial cycle and dating LIA glacial stages are novel in the context of Iberian glaciations and agree with other palaeoenvironmental studies in Iberian and in other European mountains. The limited number of boulders adequate for cosmic-ray exposure dating prevents statistical methods to be applied, and therefore highlights the need to improve geomorphological criteria in sample selection..

39 **Key words:** Sierra Nevada, cosmic-ray exposure dating, Beryllium 10, Little Ice Age, old
40 glaciations.

41

42 **1. Introduction**

43 Scientific knowledge on the glacial evolution of the Iberian mountains has greatly advanced
44 over the last decade based on the application of cosmic-ray exposure (CRE) dating on
45 moraines, erratic boulders, and polished surfaces. This method allowed inferring
46 spatiotemporal patterns of glaciation since the maximum extent of glaciers during the last
47 Pleistocene glacial cycle, with a focus on the Last Glacial Maximum (LGM) (García-Ruiz et
48 al., 2010), Oldest Dryas (Palacios et al., 2017a), and Younger Dryas (García-Ruiz et al., 2016)
49 in the main Iberian ranges such as the Cantabrian Mountains (Rodríguez-Rodríguez et al.,
50 2016), Pyrenees (Delmas, 2015), Central System (Domínguez-Villar et al., 2013), Iberian
51 System (Fernández-Fernández et al., 2017), and Sierra Nevada (Palacios et al., 2016).

52

53 The existence and distribution of landforms and deposits, their delimitation and mapping,
54 generated by glacial advances prior to the last Pleistocene glaciation are known since the first
55 geomorphological descriptions carried out in most Iberian mountains in the late nineteenth
56 and early twentieth centuries, such as in the Pyrenees (Penck, 1883; Nussbaum, 1949;
57 Barrère, 1953), Cantabrian Mountains (Penck, 1897; Obermaier, 1914), Central System
58 (Penck, 1897; Obermaier and Carandell, 1917), and Sierra Nevada (Quelle, 1908; Obermaier,
59 1916). However, the present knowledge on calendar of development of these landforms is
60 scarce, and it is based on CRE dating only in very few cases. In fact, the only CRE data
61 related to the penultimate glacial cycle were obtained with ^{21}Ne in the NW of the Iberian
62 Peninsula (Vidal-Romaní et al., 2015). These authors found glacial evidence (moraine
63 boulders and polished surfaces) of a maximum advance that occurred at 155 ± 30 ka ^{21}Ne age
64 in the Serra de Queixa, whereas polished bedrock surfaces were dated at 231 ± 48 and $131 \pm$
65 31 ka ^{21}Ne ages in the Serra da Geres. In addition, erratic blocks were dated at 113.9 ± 7.1 ka
66 ^{10}Be age in the Porma valley, Cantabrian Mountains (Rodríguez-Rodríguez et al., 2016).
67 Another erratic boulder, with a minimum ^{10}Be age of 122.2 ± 4.9 ka, was found in Ariège
68 valley (northern Pyrenees) (Delmas et al., 2011). This scarcity of CRE data from old moraines
69 may be owing to the problem of moraine degradation after being affected by intensive
70 erosion, as it has been observed even in younger moraines of other regions (Putkonen and
71 Swanson, 2003; Briner et al., 2005; Putkonen et al., 2008; Balco, 2011; Heyman et al., 2011).

72

73 The rest of the dates corresponding to glacial landforms formed during the penultimate
74 glaciation in the Iberian Peninsula were carried out by other dating methods. Villa et al.
75 (2013) were able to date with U/Th the limestone of cemented calcareous breccia in Duje
76 valley, Picos de Europa, also in the Cantabrian Mountains. This breccia deposit lies on
77 glacially abraded surfaces and is covered by moraines dated between 394 ± 50 and 276 ± 23
78 ka/ 394 ± 50 ka U/Th age. In the southern slope of the Central Pyrenees, optically stimulated
79 luminescence (OSL) dating was applied to sandy layers in fluvio-glacial terraces in two nearby
80 valleys that show the following ages: 151 ± 11 ka in the Gállego valley (Lewis et al., 2009)
81 and 263 ± 21 and 171 ± 22 ka in the Aragón valley (García-Ruiz et al., 2013).

82

83 In other Mediterranean mountains, landforms originated from glaciations prior to the Late
84 Pleistocene Glaciation (LPG) have been dated by U/Th, ^{230}Th , and OSL methods in the
85 Balkan region (Hughes et al., 2006a, Hughes et al., 2006b, Hughes et al., 2007), Dinaric Alps
86 (Hughes et al., 2011), and Apennines (Kotarba et al., 2001), but not through CRE dating. In
87 comparison to the large data set of CRE ages existing for reconstructing the LPG in the Alps,
88 only a few erratic boulders are related to the penultimate glacial cycle dated by this method
89 (Graf et al., 2015), as it is also the case of the Himalayas (Schaefer et al., 2008). On the
90 contrary, the application of ^{230}Th dates on the stalagmite record in the Alps allowed
91 determining with high resolution the last maximum ice advance of the penultimate glacial
92 cycle at 133.1 ± 0.7 and 131.9 ± 0.6 ka. These dates coincide with Heinrich stadial 11 and the
93 onset of the penultimate deglaciation or Termination II (TII), which started at 131.8 ± 0.6 ka
94 (Häuselmann et al., 2015) and reached its maximum intensity at 130.9 ± 0.9 ka (Moseley et
95 al., 2015).

96

97 However, CRE methods have never been applied to date the Little Ice Age (LIA) in the
98 Iberian Peninsula, although this cold event has been extensively documented in this region
99 (González-Trueba et al., 2008; Oliva et al., 2018a). The CRE dating showed that the LIA was
100 the most extensive Holocene glacial advance in most Iberian mountains (Palacios et al.,
101 2017b), with the exception of some cirques in the Central Pyrenees (García-Ruiz et al., 2014).
102 In fact, after pioneer studies in New Zealand (Schaefer et al., 2009) and the Andes (Licciardi
103 et al., 2009), the use of CRE dating methods to date landforms of Neoglacial and LIA age has
104 just been recently applied in the Alps (Schimmelpfennig et al., 2012, Schimmelpfennig et al.,

105 2014; Le Roy et al., 2017), Arctic (Young et al., 2015; Jomelli et al., 2016), and Central Asia
106 mountains (Li et al., 2016; Dong et al., 2017).

107

108 Radiocarbon dating on glacial, fluvioglacial, and lacustrine sediments (as well as
109 speleothems, tree rings, and historical documents) established the onset of the LIA in the
110 Iberian Peninsula at 1300 common era (CE), with the coldest conditions of ca. 2 °C below
111 present-day values at 1620–1715 CE and the last cold episodes at 1760–1800 and 1815–1835
112 CE (Oliva et al., 2018a). In Sierra Nevada there were also small glaciers during the LIA that
113 have been reconstructed based on radiocarbon dating of lake sediments (Oliva and Gómez-
114 Ortiz, 2012) and through historical documents (Gómez-Ortiz et al., 2009, Gómez-Ortiz et al.,
115 2018).

116

117 The objective of this work is to apply CRE methods to glacial landforms originated during
118 advances prior to LPG and to very young moraines as well as to review and discuss those
119 found in this research in comparison to other similar studies. Sierra Nevada constitutes a
120 unique environment to examine the use of CRE dating as a potential tool to reconstruct the
121 chronology of old and young glaciations, from the well-constrained glacial chronological
122 record of the LPG and the existence of glacial landforms originated before and after this
123 glaciation in these mountains.

124 **2. Study area**

125 Sierra Nevada is located in the SE of the Iberian Peninsula and runs SW-NE parallel to the
126 Mediterranean Sea (Fig. 1). This massif includes the highest peaks of the Iberian Peninsula:
127 Mulhacén (3478 m asl; 37°03'12"N, 3°18'41"W) and Veleta (3398 m; 37°03'02"N,
128 3°20'54"W). The area is composed of Paleozoic metamorphic rocks, mainly micaschists,
129 profoundly tectonized during the Alpine orogeny, with the presence of intensive foliation and
130 joints that makes them susceptible to intense weathering (Messerli, 1965; Sanz de Galdeano
131 and López-Garrido, 1999). Climate conditions in Sierra Nevada are characteristic of a
132 semiarid mid-latitude mountain environment, with a mean annual air temperature (MAAT) at
133 2500 m of 4.4 °C and a total precipitation of 710 mm, 40% of which falls as snow (Oliva et
134 al., 2016a). The MAAT at the top of the Veleta peak between 2002 and 2013 was 0.08 °C,
135 with a slight increase of 0.12 °C during this period (Oliva et al., 2016a).

136 The Quaternary landscape evolution of Sierra Nevada has been largely studied over the last
137 decades (Fig. 1, Table 1). From the first basic geomorphological observations of the late
138 nineteenth and early twentieth centuries to the recent multidating approaches, our
139 understanding of the glacial processes shaping the landscape of the highest lands has
140 significantly improved (i.e., Gómez-Ortiz et al., 2015).

141 During the 1960s and 1970s several authors suggested that the existence of older glaciations
142 occurred prior to the LPG, formerly known as the Riss glaciation in Alpine terminology.
143 Their observations focused on the existence of highly eroded moraine remnants and
144 fluvioglacial deposits at very low altitudes, mainly in the southern slope of the massif
145 (Hempel, 1960; Messerli, 1965; Lhenaff, 1977; Sánchez-Gómez, 1990). However, this
146 hypothesis has not been validated until now though direct dating methods.

147

148 By contrast, the LPG has been recently well-constrained with CRE techniques. Gómez-Ortiz
149 et al. (2012) and Palacios et al. (2016) reported that two major glacial advances occurred at
150 ca. 30–32 and 19–20 ka. The second stage of glacial advance occurred in phase to the global
151 Last Glacial Maximum and reached almost the same extent than during the previous glacial
152 stage. Subsequently, temperature increase recorded worldwide (Clark et al., 2009) favored a
153 massive deglaciation of the massif, although two stages of glacial expansion occurred during
154 the transition toward the Holocene, namely during the Oldest Dryas and Younger Dryas.
155 During these phases, glaciers expanded significantly and flowed downvalleys several
156 kilometers, in southern and northern slopes (Gómez-Ortiz et al., 2012; Palacios et al., 2016).

157

158 The Holocene was characterized by warmer temperatures that promoted the gradual melting
159 of these glaciers during the early Holocene (Palacios et al., 2016). The formerly glaciated
160 environments were then subjected to paraglacial processes, which favored the development of
161 rock glaciers associated to permafrost conditions (Oliva et al., 2016b, Oliva et al., 2018b).
162 These permafrost-related landforms finally stabilized at 6–7 ka during the Holocene Climate
163 Optimum, when warmer conditions must have conditioned permafrost thawing (Palacios et
164 al., 2016).

165

166 Climate variability intensified in the Northern Hemisphere during the late Holocene
167 (Mayewski et al., 2004). In Sierra Nevada, this was reflected in alternating cold and warm

168 phases (Oliva et al., 2014). During the coldest stages, some glacier spots developed in the
169 highest northern cirques, such as in the Mulhacén cirque, where lake sediments from La
170 Mosca Lake revealed evidence of the existence of a glacier within the cirque at ca. 2.8–2.7
171 and 1.4–1.2 ka cal BP and during the LIA between 1440 and 1710 CE (Oliva and Gómez-
172 Ortiz, 2012). For the last cold stage, historical documents also showed evidence of the
173 prevailing colder conditions in Sierra Nevada, with descriptions sketches and pictures of
174 glaciers and permanent snow fields at the foot of the highest peaks of the westernmost
175 mountains of Sierra Nevada between the late seventeenth and the early twentieth centuries.
176 The last glacier was located in the Veleta cirque and constituted the southernmost glacier in
177 Europe during the LIA, although it finally melted away during the mid-twentieth century
178 (Gómez-Ortiz et al., 2001, Gómez-Ortiz et al., 2009, Gómez-Ortiz et al., 2018; Oliva and
179 Gómez-Ortiz, 2012; Oliva et al., 2018a). However, no CRE dating is yet available from the
180 different LIA glacial stages in Sierra Nevada.

181

182

183 **3. Methods**

184 *3.1 Sampling strategy*

185 With the aim of dating glacial landforms formed potentially before the LPG, this research
186 focuses on the Naute valley, a steep mountain ravine located on the southern slope of the
187 Mulhacén peak (Fig. 1, Fig. 2, Fig. 3, Fig. 4). The valley results from the confluence of two
188 glacial headwaters — Río Seco and Mulhacén valleys — converging at 2300 m of altitude. In
189 previous studies these two high mountain valleys were selected to reconstruct glacial stages of
190 the LPG using CRE methods (Gómez-Ortiz et al., 2012; Palacios et al., 2016), but landforms
191 distributed in the lower Naute valley affected by intense erosion processes were not
192 examined. In Río Seco and Mulhacén valleys landforms from the LPG have been already
193 dated using the in situ cosmogenic nuclide ^{36}Cl , such as the front of a fossil rock glacier (9.0
194 ± 0.8 ka) resting on polished rock surfaces (11.9 ± 1.1 ka) in the Río Seco cirque, as well as of
195 a fossil rock glacier (6.3 ± 0.5 and 13.1 ± 1.3 ka) resting on a polished rock surface ($11.2 \pm$
196 1.1 and 12.7 ± 1.2 ka) at the head of the Mulhacén valley (Palacios et al., 2016). Just before
197 their confluence, both valleys include a sequence of moraine ridges, where five blocks were
198 dated using ^{36}Cl (29.9 ± 2.6 , 16.9 ± 1.8 , 14.9 ± 1.5 , 13.7 ± 1.3 , and 11.7 ± 1.6 ka; Palacios et
199 al., 2016). These authors interpreted these ridges as a polygenic moraine formed by

200 successive advances generated from just before the LGM to the Oldest Dryas. In the Naute
201 valley, immediately below the confluence of these two tributaries, a lateral moraine is
202 preserved in the left sector of the valley at elevations between 2150 and 2300 m. This moraine
203 is placed below one of the ridges where a boulder was dated at 29.9 ± 2.6 ka (Palacios et al.,
204 2016), and therefore must be older in age, possibly from a previous glaciation. For this
205 reason, this lower moraine was considered as one of the objectives of the work in order to
206 apply the CRE dating method.

207 To date glacial oscillations during the LIA by CRE dating methods, we focused our research
208 on the Veleta cirque, located under the northern wall of Veleta peak (Fig. 1, Fig. 5, Fig. 6).
209 Historical documents provided abundant evidence of a glacier existing at the floor of this
210 cirque during the LIA until the first half of the twentieth century (Gómez-Ortiz et al., 2001,
211 Gómez-Ortiz et al., 2009, Gómez-Ortiz et al., 2018). The cirque is closed by a large frontal
212 moraine whose age is unknown, although because of its large size it has been considered to
213 have formed during the Last Termination or during Holocene advances (Oliva et al., 2014,
214 Oliva et al., 2018a).

215 .

216 *3.2 Sampling and analytical procedures*

217 even samples were collected from moraine boulders using a hammer and a chisel: four located
218 in Naute valley (NAUT-1, -2, -3, -4) between 2164 and 2263 m, and three in the Veleta cirque
219 (SN-11-1, -2, -3) between 3061 and 3095 m (Table 1). All samples were taken from the flat-
220 topped surfaces of boulders >1 m high, located on the crests of the moraines. Field data and
221 sample characteristics of the seven samples are listed in Table 2.

222

223 Physical and chemical sample preparation and beryllium measurements were carried out at
224 the Centre Européen de Recherche et d'Enseignement des Géosciences de l'Environnement
225 (CEREGE, France). Lichen, moss and other organic fragments were removed from the
226 samples with a brush. Samples were crushed with a roller grinder and sieved to retrieve the
227 grain size fraction 0.25–1 mm. All samples are quartz-bearing micaschists, and therefore we
228 selected in situ-produced cosmogenic ^{10}Be dating to determine the ages of the moraines. To
229 isolate the quartz from the bulk rock, the samples went repeatedly through a magnetic
230 separator (Frantz LB-1) until all magnetic minerals were discarded. Subsequently, the
231 nonmagnetic fraction experienced successive chemical attacks with a mixture of concentrated

232 hydrochloric (HCl) and hexafluorosilic (H₂SiF₆) acids to dissolve the remaining nonquartz
233 minerals. The residual impurities were dissolved when the sample grains were
234 decontaminated from meteoric ¹⁰Be by three successive partial dissolutions with concentrated
235 hydrofluoric acid (HF).

236

237 The subsequent beryllium extraction protocol is adapted from Brown et al. (1991) and
238 Merchel and Herpers (1999) chemical procedures. The samples yielded between 18 and 32 g
239 of purified quartz (Table 3). About 100 µl of a ⁹Be carrier solution with a ⁹Be concentration
240 of 3025 µg/g prepared in-house from a phenakite crystal was added to the quartz before it was
241 dissolved in HF. A chemistry blank was prepared along with the seven samples. Following
242 evaporation of the resulting solution, the samples were recovered in a hydrochloric acid
243 solution and beryllium was precipitated to Be(OH)₂ with ammonia before and after elution
244 through an anionic exchange column (Dowex 1X8) to remove iron. Following the methods
245 described by Merchel and Herpers (1999), a cationic exchange column (Dowex 50WX8) was
246 used to remove boron and to separate the Be from other elements. Beryllium was precipitated
247 with ammonia to Be(OH)₂ and the resulting precipitate was oxidized to BeO at 700 °C. Then,
248 this final BeO was mixed with niobium powder and loaded on cathodes for analysis of the
249 ¹⁰Be/⁹Be ratios at the French national Accelerator Mass Spectrometry (AMS) facility ASTER
250 (CEREGE) (Arnold et al., 2010). The measurements were calibrated against in-house
251 standard STD-11 using an assigned ¹⁰Be/⁹Be ratio of $1.191 (\pm 0.013) \times 10^{-11}$ (Braucher et al.,
252 2015). Sample ¹⁰Be/⁹Be ratios were corrected for the chemical blank background by
253 subtracting the measured chemistry blank ¹⁰Be/⁹Be ratio (Table 3). Analytical 1σ uncertainties
254 include uncertainties in AMS counting statistics, the uncertainty in the standard ¹⁰Be/⁹Be
255 ratio, an external AMS error of 0.5% (Arnold et al., 2010), and the uncertainty in the chemical
256 blank measurement. A ¹⁰Be half-life of $1.387 (\pm 0.01) \times 10^6$ years was used (Chmeleff et al.,
257 2010; Korschinek et al., 2010).

258

259 *3.3 Age calculation (¹⁰Be age computation)*

260 We calculated the ¹⁰Be CRE ages with the online CREP exposure age calculator (Martin et
261 al., 2017; <http://crep.crpq.cnrs-nancy.fr>), using the LSD scaling model (Lifton et al., 2014)
262 that is similar to other previous empirical models (Borchers et al., 2016). Age calculations
263 have also considered the ERA40 atmospheric model (Uppala et al., 2005) and LSD

264 Framework geomagnetic database (Lifton et al., 2014). As it is not regional production rate
265 available, we used the worldwide mean ^{10}Be spallation production rate of 3.99 ± 0.22 atoms
266 $\text{g}^{-1} \text{y}^{-1}$, as calibrated in the ICE-D production rate database linked to CREP. For all samples,
267 a rock density of 2.7 g cm^{-3} was considered, and the topographic shielding factor of each
268 sample was calculated (Table 2).

269

270 In the moraine of the Veleta cirque, from where samples SN-11-1, SN-11-2, and SN-11-3
271 were extracted, snow cover remains for a large part of the year (Gómez-Ortiz et al., 2009);
272 and there is information on the properties and persistence of the snow cover for the last
273 decades (Herrero and Polo, 2016), which allows considering snow shielding effect for these
274 samples. The correction for snow cover has been calculated applying the equation by Gosse
275 and Phillips (2001), including local parameters (Herrero and Polo, 2016), and considering a
276 snow thickness on moraine boulders ranging from 30 to 130 cm during 8 months per year. We
277 used an average value of snow density of 0.3 g cm^{-3} , considering that the attenuation length
278 for fast neutrons in snow is 109 g cm^{-2} (Zweck et al., 2013; Delunel et al., 2014). Taking into
279 account these characteristics, we applied a snow shielding factor of 0.865 to the affected
280 samples. We have not applied a snow shielding factor to the old samples of Naute moraine, as
281 the coverage of snow in this area is much shorter and data are not available for a period
282 encompassing several tens of thousands of years.

283

284 Table 3 also includes the ages determined with version 3 of the online exposure age calculator
285 formerly known as the CRONUS-Earth online exposure age calculator (Balco et al., 2008;
286 Balco, 2018). These ages have been calculated using the default calibration data set based on
287 the ICE-D calibration database and the time-dependent LSDn scaling method (Lifton et al.,
288 2014). The average difference between the ages calculated with the CREP exposure age
289 calculator and those calculated with the CRONUS-Earth (v.3) is 2.4% for NAUT samples and
290 6.9% for SN-11 samples.

291

292 The ages of the SN-11 samples calculated with the snow shielding factor turn out to be
293 between 8.8 and 16.9% higher in the online CREP exposure age calculator (Martin et al.,
294 2017) and between 7.9 and 16.9% higher in the online exposure age calculator formerly
295 known as the CRONUS-Earth online exposure age calculator v. 3 (Balco et al., 2008; Balco,

296 2018). The CREP exposure calculator ages and the analytical uncertainties are used in the text
297 and in the figures of this manuscript and, only in the case of the Veleta cirque, with the
298 application of the snow shielding factor (Table 3).

299 300 **4. Results**

301 302 *4.1 Geomorphological analysis of the Naute moraine and CRE dating results*

303 To select the most suitable boulders for CRE dating, an exhaustive survey was carried out of
304 the Naute moraine system, formed potentially during a glacial advance before the LPG. The
305 results of this survey are expressed in a detailed geomorphological map (Fig. 2). The moraine
306 is located on the left bank of the Naute river (west) and the Peñón Grande (east) between
307 2150 and 2300 m. It has been intensely reshaped by postglacial environmental dynamics (i)
308 fluvial processes driven by steep streams have deeply eroded the moraine ridge, and (ii)
309 gullyng processes have washed away a large part of its abundant fine sediments. As a result,
310 a large number of boulders — highly weathered and fragmented caused by their abundant
311 foliation planes and joints — are found on the surface of the moraine (Fig. 3).

312
313 To sample the boulders that best represent the original surface of the moraine, we focused on
314 blocks that (i) present aligned upper edges indicating the maximum possible height of the
315 original moraine crest, and (ii) are inserted into the moraine, which ensures that they have not
316 been mobilized since the moraine stabilization. Using these criteria, only four boulders were
317 found with these characteristics (NAUT-1, -2, -3, and -4) (Fig. 4). The rest of the boulders
318 showed traces of having been affected by erosion and exhumed.

319
320 Samples from two boulders showed a similar age: NAUT-1 yielded 134.8 ± 6.2 ka and
321 NAUT-4 129.2 ± 5.3 ka. The other two boulders showed different and substantially younger
322 ages: NAUT-2 obtained 53.7 ± 2.9 ka and NAUT-4 43.8 ± 2.3 ka (Fig. 2, Fig. 3; Table 3).

323 324 *4.2 Geomorphological analysis of the Veleta cirque moraine and CRE dating results*

325 The Veleta cirque moraine is composed of a large proportion of fine sediments with a few
326 large blocks on the surface, possibly deposited by different glacial advances. Within the area
327 enclosed by this moraine, there are large rock avalanche deposits fallen from the Veleta

328 northern wall and a small rock-glacier where buried glacial ice is still preserved under the
329 debris cover, though it presents evidence of accelerated degradation. On the other hand, the
330 fine sediments of the moraine, resulting from heavily weathered micaschists, are being
331 washed away by runoff activity. This process is particularly intense in late spring and early
332 summer with snowmelt runoff and in autumn with torrential rain events. The outer slope of
333 the moraine is being reshaped by frequent debris flows, which also facilitate the outcropping
334 and exhumation of large blocks buried in the moraine. In addition, some metric-size boulders
335 are present fallen from the Veleta north wall and deposited on the surface of the moraine,
336 normally during the winter season when snow completely fills the floor of the cirque. These
337 rocks slide on the snow and ice surface and stabilize on the moraine ridge. On the inner slope
338 of the moraine, accumulations are found of superimposed till blocks of different sizes. In
339 contrast to the main moraine, these till deposits include a smaller proportion of fine particles
340 but conserve a recent glacial imprint as revealed by the presence of flutes. This deposit
341 exceeds the limits of the main moraine on the eastern side of the cirque, where the moraine
342 height is lower (Fig. 5, Fig. 6). Evidence from historical sources confirms that the distribution
343 of this recent till deposit coincides with the surface covered by the glacier during the
344 nineteenth century (Fig. 7; Bide, 1893; Gómez-Ortiz et al., 2018).

345 Before selecting boulders for CRE dating, a very detailed geomorphological survey of the
346 large polygenic moraine system enclosing the Veleta cirque was conducted. A detailed
347 geomorphological map shows the main landforms observed in the cirque (Fig. 5). Schematic
348 transects show the sectors of the moraine affected by current or recent post-glacial processes
349 and derived landforms (Fig. 6). Considering these geomorphological characteristics, we
350 avoided collecting samples from blocks corresponding to very recently deglaciated surfaces
351 because (i) we already know that this area was occupied by the glacier during the nineteenth
352 century and (ii) they can result from the frequent rockfalls from the headwall. As a result,
353 only three CRE datable boulders were found anchored in the moraine and protruding
354 sufficiently not to have been covered by other sediments since their deposition (SN11-1, -2,
355 and -3) (Fig. 7). Thus, they cannot have fallen from the headwall or have been covered by the
356 glacier in a subsequent readvance.

357

358 Sample SN-11-1, collected on the top of the moraine, very close to twentieth century till
359 deposits, obtained an age of 340 ± 120 years, corresponding to the year 1675 ± 120 CE.
360 Sample SN-11-2, collected on the crest of the lobate ridge of the moraine, yielded an age of

361 720 ± 270 years (1295 ± 270 CE). Finally, sample SN-11-3, collected on the top of the lateral
362 moraine ridge in the western part of the cirque, obtained an age of 400 ± 110 years (1615 ±
363 110 CE) (Fig. 5, Fig. 6; Table 3).

364

365 **5. Discussion**

366 *5.1 The existence of a glaciation prior to the last Pleistocene cycle in Sierra Nevada*

367 The results of ¹⁰Be CRE dating yielded ages of ca. 130–135 ka for the two samples collected
368 from boulders in the moraine downslope of the maximum ice advance in Sierra Nevada
369 (Gómez-Ortiz et al., 2012). These ages suggest that the moraine may have been formed
370 during the last maximum advance of the penultimate glaciation. The other two boulders of the
371 same moraine yield much younger ages, around 45–50 ka (Table 3), and thus can be
372 considered outliers. On the other hand, none of the studies in Sierra Nevada reported boulder
373 samples affected by nuclide inheritance (Gómez-Ortiz et al., 2012, Gómez-Ortiz et al., 2015;
374 Palacios et al., 2016), and therefore this problem can be rejected in the older samples (Fig. 8).

375

376 In Sierra Nevada, no polished and striated bedrock surfaces were found in this study between
377 the moraine ridges dated ca. 130–135 ka and those dated in previous studies between ca. 30
378 and 15 ka (Palacios et al., 2016), i.e., rock surfaces that may have been ice-covered prior to
379 the LPG. All the polished bedrock surfaces dated in Sierra Nevada — in many cases with
380 abundant and well-preserved fresh striations — correspond to the deglaciation process during
381 Termination I (Gómez-Ortiz et al., 2012; Palacios et al., 2016). The fact that moraines from a
382 previous glaciation are found in close proximity to those of the last glacial cycle without
383 intermediate polished surfaces led us to apply CRE dating only on boulder surfaces in the old
384 moraines. Recent research in mountain areas, such as in the Alps (e.g., Chenet et al., 2016),
385 the Andes (e.g., Martini et al., 2017), Central Asia (e.g., Batbaatar et al., 2018), and the
386 Iberian Peninsula (Rodríguez-Rodríguez et al., 2016, Rodríguez-Rodríguez et al., 2017),
387 suggest using a minimum of five boulders from the same moraine to ensure obtaining the
388 formation age of the landform. Thus, where most of the results show dispersed data, the ages
389 are rejected, whereas an average value is obtained with the rest, interpreted as indicative of
390 the development age of the moraine. In this study we attempted to apply this five-boulder
391 criterion, but this proved impossible because of the deterioration of the moraine since the
392 penultimate glaciation, as the landform has been intensely eroded and few potentially datable

393 boulders are available. The headwaters of several torrents have eroded the ridge summit,
394 leaving the moraine covered with loose blocks. When trying to reconstruct the original
395 surface of the moraine, we found only four embedded boulders (Fig. 4).

396

397 The CRE ages of ca. 130–135 ka obtained in Sierra Nevada coincide with the last cold period
398 of the penultimate glaciation in the Alps as suggested by ^{230}Th dating on speleothems that
399 obtained 131.8 ± 0.6 ka (Häuselmann et al., 2015; Moseley et al., 2015). This stage occurred
400 during Heinrich stadial 11 (Oppo et al., 2006), parallel to the end of the last cold phase in the
401 Alboran sea surface temperatures (near Sierra Nevada, in SE Iberia) before the beginning of
402 TII (Martrat et al., 2014; Jiménez-Amat and Zahn, 2015). By that time, the onset of warmer
403 temperatures favored the end of the last cold period of the penultimate glaciation in Greenland
404 (Grant et al., 2012) as well as the initial melting of large ice sheets at high latitudes in the
405 Northern Hemisphere (Govin et al., 2015), leading to sea level rise (Rohling et al., 2017).

406

407 The formation and stabilization of the Naute moraine may therefore be contemporary with
408 what is called the Penultimate Glacial Maximum (PGM), a parallel concept to the LGM for
409 the last glaciation, centered on 140 ka (Rohling et al., 2017). The deposition age of the Naute
410 moraine is similar to the stabilization age of the maximum glacial advance moraine in the
411 Sierra de Queixa dated 155 ± 30 ka ^{21}Ne (Vidal-Romaní et al., 2015). The erratic boulder
412 dated 113.9 ± 7.1 ka ^{10}Be in the Porma Valley, Cantabrian Mountains, could also be related to
413 the deglaciation process following the PGM (Rodríguez-Rodríguez et al., 2016). The same is
414 true for the erratic boulder in the Ariège valley (northern Pyrenees), with a minimum ^{10}Be age
415 of 122.2 ± 4.9 ka (Delmas et al., 2011), slightly after the formation of the Naute moraine from
416 the penultimate glaciation. In addition, fluvio-glacial terraces deposited simultaneously on the
417 outermost moraine ridges in the Gállego and Aragón valleys, Central Pyrenees, were dated
418 151 ± 11 ka and 171 ± 22 ka OSL ages by other authors (Lewis et al., 2009; García-Ruiz et
419 al., 2013), and therefore may represent the same glacial stage that led to the formation of the
420 Naute moraine. However, despite some evidence suggesting glacial activity between ca. 170
421 and 130 ka in the Iberian Peninsula, the small number of reliably dated terrestrial records
422 makes it difficult to infer spatiotemporal glaciation patterns occurring before the LPG. This is
423 a common issue in other Mediterranean mountains where glacial advances concurrent with
424 the development of the Naute moraine have only been dated through the application of U

425 series to secondary carbonates in cemented moraines in mountains in Greece (Hughes et al.,
426 2006a, Hughes et al., 2006b, Hughes et al., 2007), Dinaric Alps (Hughes et al., 2007) and the
427 Gran Sasso Massif, Central Italy (Kotarba et al., 2001).

428

429 A case study similar to the Naute moraine was undertaken in the Jura mountains, where
430 several erratic boulders were found scattered beyond the Alpine glaciers LGM extent. In this
431 case the oldest ages were very similar to those of the Naute moraine (129.7 ± 4.7 ka, $143.2 \pm$
432 8.2 , and 144.0 ± 5.3 ka ^{10}Be ages), although just as in Sierra Nevada, other boulders in the
433 same deposit yielded younger ages (Graf et al., 2015). After their experience applying CRE
434 dating to several moraine boulders in the Himalayas, Schaefer et al. (2008) highlighted the
435 role of erosion processes in the exhumation of blocks in old moraines. As a result, these
436 authors claim that the oldest, not the average age of the moraine boulders reflects the best
437 minimum age estimation for the moraine deposition, in agreement with previous authors
438 (Hallet and Putkonen, 1994). Briner et al. (2005) proposed the same criterion after analyzing a
439 large number of moraine boulders, presumably from the penultimate glaciation in Alaska, and
440 found many problems of boulder exhumations. These authors also defend an oldest-age
441 method to provide the closest approximation to the moraine stabilization age. Similarly, and
442 from previous experience, Balco (2011) defined the most appropriate strategies when
443 applying CRE methods to degraded moraines, which have been considered in this study.

444

445 From our results we conclude that the difficulty of finding suitable boulders to apply CRE
446 methods in moraines from old glaciations should not discourage their use. Dating an adequate
447 number of samples is very difficult for moraines prior to the LPG, where erosive activity over
448 time may have exhumed most of the boulders. In the case of Sierra Nevada, the profoundly
449 tectonized micaschists have been intensely affected by erosion, which makes it more difficult
450 to find suitable boulders for dating (Gómez-Ortiz et al., 2015). Nevertheless, Sierra Nevada is
451 an arid mountain, where ice margin deposits preserved on gentle slopes make it possible to
452 obtain a well-dated glacial sequence using the CRE method (Pallàs et al., 2010) (Fig. 9).

453

454 *5.2 Chronology of the formation of LIA moraines in Sierra Nevada*

455 It was also difficult to find suitable blocks for CRE dating in young moraines in Sierra
456 Nevada, as they are found in very dynamic geomorphological environments on steep slopes of

457 200 m high (up to 300 m from the peaks to the cirque floors) with thick and long-lasting snow
458 cover. On the one hand, these factors accelerate erosion processes in the moraine, often
459 reducing its size, vertical development, and the slope gradient; on the other hand, they favor
460 new sediment deposits on the moraine surface. These difficulties are also observed in most
461 Mediterranean mountains, where LIA glaciers did not exceed the limits of the cirques and
462 developed at the foot of vertical walls (Hughes, 2014, Hughes, 2018). Previous studies
463 adverted to the polygenic character of the Veleta cirque moraine, related to the effects of an
464 Early Holocene glacial advance that pushed downslope, accumulating rockfall deposits from
465 the end of TI (Gómez-Ortiz et al., 2012; Palacios et al., 2016). According to these authors,
466 this moraine system functioned as a barrier to subsequent advances, which accumulated
467 sediments on the inner slope, generating a single, large, polygenic moraine.

468

469 Using the criteria defined in Methods above, we obtained three coherent results to infer the
470 age of the moraine, although dates were obtained for only three boulders. The high
471 uncertainty levels in our results prevent us from relating the age of each sample to LIA
472 periods (Oliva et al., 2018a). These ages could be related to advances from 1300 to 1610/1680
473 CE, contemporary with the LIA maximum glacial advance in the Alps (Holzhauser et al.,
474 2005) and the lowest temperatures in the Iberian Peninsula during the Maunder Minimum
475 (Oliva et al., 2018a).

476

477 The inner slope of the moraine therefore developed during the LIA, although the large
478 moraine system may also include boulders from other Neoglacial glacial stages, such as those
479 detected in the Mulhacén cirque where lake sediments show evidence of a glacier within the
480 cirque ca. 2.8–2.7 and 1.4–1.2 ka cal BP and LIA (Oliva and Gómez-Ortiz, 2012).

481

482 New results obtained in Mount Olympus (Greece) found similar results using ^{36}Cl dating
483 method, with one sample of 0.64 ± 0.08 ka (Styllas et al., 2018). However, most studies
484 focusing on the use of CRE dating methods for dating LIA glacial advances have been carried
485 out in the Alps. In the Western Alps, LIA moraines are very close to Neoglacial (Late
486 Holocene) moraines and may even overlap them (Schimmelpfennig et al., 2012,
487 Schimmelpfennig et al., 2014); as in the Veleta cirque, LIA glacial advances have left a single
488 polygenic ridge with dated boulders showing a wide range of dates: 1430 ± 32 , 1534 ± 28 ,

489 and 1829 ± 11 CE ^{10}Be (ages not considering snow shielding) (Schimmelpfennig et al., 2012).
490 These dates could correspond to the three most important glacial advances in the Alps during
491 the LIA: 1300–1380, 1600–1670, and 1800–1860 CE (Holzhauser et al., 2005), similar to
492 those detected in the Pyrenees (Oliva et al., 2018a). Recent studies in the central Alps showed
493 a similar pattern when dating LIA moraine boulders — also very close to or overlapping
494 Neoglacial moraines — forming single polygenic crests (Schimmelpfennig et al., 2014).
495 These authors dated 14 boulders in a single LIA moraine ridge using ^{10}Be , assuming
496 negligible inheritance for the boulders, and reported dates ranging from 1430 to 1870 CE
497 (Schimmelpfennig et al., 2014). Le Roy et al. (2017) dated Neoglacial moraines near several
498 present-day glacier fronts in the French Alps and obtained ^{10}Be ages that are all consistent
499 with the late Holocene period (~ 4 – 1 ka), but do not follow the logical chronostratigraphic
500 moraine sequence. These authors highlighted the problems related to surface exhumation and
501 erosion of many moraine ridges.

502

503 Some attempts to date the moraines of the LIA have also been made in other high mountain
504 ranges such as the Cordillera Vilcabamba, central Andes where Licciardi et al. (2009) applied
505 ^{10}Be CRE dating for the first time. Samples collected from a single massive ridge showed
506 average dates ranging from 1740 ± 30 to 1810 ± 20 CE, though samples of old dates were
507 rejected as being outliers. They suggest that the moraines found in several valleys of this
508 massif may be the result of the successive accumulation of multiple late Holocene glacial
509 expansions.

510

511 Schaefer et al. (2009) also applied ^{10}Be dating in the Mueller valley in the Southern Alps,
512 New Zealand, to date many LIA moraine boulders. The geomorphological setting is similar to
513 that described in the European Alps, with a single LIA moraine system overlapping with
514 various late Holocene moraine ridges. Boulders taken from different sites in the LIA moraine
515 reported dates of 1350 ± 60 , 1600 ± 50 , 1780 ± 40 , and 1820 ± 20 CE ^{10}Be , which coincide
516 with cold periods inferred from tree-ring data. Li et al. (2016) applied ^{10}Be dating in LIA
517 moraines in Tian Shan mountains, central Asia; interestingly, they found that boulders in
518 moraines from glaciers smaller than 1.0 km² show very old ages because of nuclide
519 inheritance. The same problems were found in the small Veleta cirque, where the wall is very
520 close to the moraine, but we discarded this possibility following the results obtained in this

521 present study and those from other previous work in Sierra Nevada, because of the intense,
522 continuous rock fall activity on its wall. For larger glaciers, Li et al. (2016) found evidence of
523 a major advance ca. 1600 ± 100 CE ^{10}Be and also a remarkable early LIA glacial expansion at
524 ca. 1480 ± 55 CE ^{10}Be . Dong et al. (2017) dated two different LIA ridges in Tibetan
525 mountains with ^{10}Be , collecting four samples from each ridge. Their results provide a range of
526 1480 ± 139 to 1975 ± 31 CE ^{10}Be dates for the entire moraine system, very similar to the
527 Veleta cirque results. Young et al. (2015) dated several stable moraine ridges close to the
528 current alpine glacier snouts in Baffin Island and western Greenland and found a logical
529 geomorphological order from 1040 ± 40 to 1700 ± 40 CE ^{10}Be ages with a retreat from 1750
530 CE and proposed a recent glacial maximum during the Medieval Climate Anomaly, when
531 glaciers in Europe receded. Jomelli et al. (2016) applied in situ cosmogenic ^{36}Cl to date three
532 moraines in Lyngmarksbræen glacier (west Greenland), which were deposited during the last
533 millennium in a relatively flat area not constrained by topography, the opposite that happens
534 in Alpine environments. These authors obtained advances with average ages of 1200 ± 130 ,
535 1450 ± 90 , and 1720 ± 60 CE for the most external and internal ridges. The most recent
536 advances coincide with glacial expansion in European high mountains, but the older moraine
537 may have developed during the Medieval Climate Anomaly (Jomelli et al., 2016). In the sub-
538 Antarctic Kerguelen Islands, Jomelli et al. (2017) recently dated a moraine with ^{36}Cl that is
539 probably related to the LIA.

540

541 From a review of the available literature on LIA moraine dating in different mountain ranges,
542 we concluded that, in general, a first glacial advance was recorded during the fourteenth
543 century, with the major ice expansion during the seventeenth century and a last minor
544 readvance at the beginning of the nineteenth century.

545

546 Owing to the short exposure duration of our samples, their ^{10}Be ages have 25–30% analytical
547 uncertainties. These uncertainties are higher than the age difference of 10–16% obtained when
548 a snow shielding correction factor is applied, as was proposed in early studies (Benson et al.,
549 2004; Schildgen et al., 2005). In any case, this study defends the importance of considering
550 the snow cover of the moraines studied, when applying CRE methods and when discovering
551 the degree of degradation to which they have been subjected. The location of the glacial fronts
552 during the LIA suggests MAAT ca. $1\text{ }^{\circ}\text{C}$ below current values in mid-latitude mountain

553 regions (Oliva et al., 2018a), determining longer snow cover duration in the highlands of the
554 glaciated massifs and strengthening shielding from cosmogenic radiation (Fig. 7). Long-
555 lasting snow cover also intensifies nivation-related erosion processes and can intensely affect
556 glacial deposits (Christiansen, 1998; Palacios et al., 2003).

557

558 On the other hand, as we observed in the Veleta cirque and as described in several papers
559 cited above, the intensity of the paraglacial processes (Ballantyne, 2002; Oliva and Ruiz-
560 Fernández, 2015) in the current deglaciation phase from the LIA advance is very high, with a
561 complex superposition of processes that destroy the glacial landforms and accumulate new
562 deposits on them. Therefore, as we saw, it was difficult to find suitable boulders for CRE
563 dating in the Veleta cirque, despite its young age. In this context, the long-term study and
564 monitoring of the dynamics of the LIA moraines in full paraglacial phase, should serve to
565 evaluate the time needed for moraine stabilization once it has been abandoned by glacial
566 retreat (Zreda and Phillips, 1994, Zreda and Phillips, 1995; Putkonen and O'Neal, 2006;
567 Putkonen et al., 2008; Heyman et al., 2011) (Fig. 9).

568

569 In areas where LIA moraines are constrained by steep slopes, as is very common in Alpine
570 systems, at least two (or possibly three) major LIA advances overlap, in some cases with
571 other Neoglacial moraines, to form a single polygenic ridge as occurred in the Veleta cirque.
572 This was highlighted in the mid-1980s with the concepts of oblitative overlap (Gibbons et
573 al., 1984) and distal-flank accretion (Osborn, 1986) applied to Neoglacial advances.

574 **5. Conclusions**

575 The results of this research supply evidence of the very few CRE dates available on very old
576 glacial landforms formed before the last Pleistocene glaciation and also the few very young
577 dates available, such as those derived from LIA glaciation, compared with the large number
578 of dates obtained for Termination I. For both periods, the selection of boulders suitable for
579 sampling was the critical issue in Sierra Nevada, caused by intense postglacial environmental
580 dynamics such as occurred during the last major deglaciation, which impedes the collection of
581 a statistically significant number of samples. In the case of the LIA, moraines are still
582 undergoing intense paraglacial readjustment to the new geomorphological setting, with very
583 intense erosion and sediment redistribution of the unconsolidated moraine particles. In
584 addition, moraine accretion processes during successive glacial advances are detected from

585 geomorphological observations. In moraines originating during the penultimate glaciation, the
586 long time lapse since their formation has facilitated intense erosion and reshaping of these
587 landforms, with very few stable blocks remaining on the surface since their original
588 stabilization. Periods of certain geomorphic stability may have occurred between the very old
589 and the younger glaciations (Fig. 9).

590

591 The results of this research have enabled us to verify the existence of moraines deposited in
592 glaciations prior to LPG in Sierra Nevada, as many of the first geomorphologists who studied
593 the glacial landforms of these mountains suggested, and to confirm similar CRE and OSL
594 ages established in other mountains in the northern Iberian Peninsula.

595

596 In addition, CRE dates are obtained for the first time in the Iberian Peninsula in LIA
597 moraines, with results suggesting that maximum advances during this period occurred
598 between the fourteenth and seventeenth centuries, as is the case in many other mountains
599 where the same method has been applied.

600

601 **Acknowledgments**

602 The research was carried out within the MOUNTAIN WARMING project CGL2015-65813-
603 R, Government of Spain. Marc Oliva is grateful for the support of the Ramón y Cajal research
604 program (RYC-2015-17597) and the ANTALP research group (Antarctic, Arctic, Alpine
605 Environments, 2017-SGR-1102). The ^{10}Be measurements were performed at the ASTER
606 AMS National facility (CEREGE, Aix en Provence) which is supported by the INSU/CNRS,
607 the ANR through the “Projets thématiques d'excellence” program for the “Equipements
608 d'excellence” ASTER-CEREGE action and IRD. The authors express their deep gratitude to
609 Dr. Magali Delmas and two anonymous reviewers whose detailed and interesting suggestions
610 have helped to improve our manuscript. We are also grateful to Editor Richard Marston for
611 his useful suggestions and patient editing.

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934

Table 1. Current knowledge of Quaternary glacial stages in Sierra Nevada.

Stage	Chronology	Processes and landforms
Pre-Last Glaciation	unknown	Possible existence of eroded moraines from glaciations occurred before the Last Glaciation as well as glacio-fluvial sediments distributed at lower elevations than moraines formed during the Last Glaciation (Hempel, 1960; Messerli, 1965; Lhenaff, 1977; Sánchez-Gómez, 1990).
Last Glaciation (Maximum Ice Extent; MIE)	Two glacial advances at ca. 30-32 ka and 19-20 ka	Development of alpine valleys at elevations between 2000 and 2500 m in northern and southern slopes, respectively (Gómez-Ortiz et al., 2012, Gómez-Ortiz et al., 2013, Gómez-Ortiz et al., 2015; Oliva et al., 2014; Palacios et al., 2016). The glaciated environment during the MIE encompassed 105 km ² (Palma et al., 2017).
Deglaciation	Two glacial advances during the Oldest Dryas (OD) at ca. 17 ka and Younger Dryas (YD) at ca. 12-13 ka	Glacial retreat at ca. 19 ka followed by two phases of glacial development until the Holocene, mainly in the highest valleys from the northern slope of the massif (Gómez-Ortiz et al., 2012, Gómez-Ortiz et al., 2013; Oliva and Gómez-Ortiz, 2012; Oliva et al., 2014; Palacios et al., 2016).
Holocene	Glacial retreat until the Early Holocene at ca. 9-10 ka Three stages with formation of glaciers occurred during the Late Holocene at ca. 2.8-2.7, 1.4-1.2 ka cal BP and LIA	The glaciers formed during the YD finally melted and paraglacial activity favoured the development of rock glaciers in these areas (Gómez-Ortiz et al., 2012, Gómez-Ortiz et al., 2013; Oliva et al., 2016b; Palacios et al., 2016). Lake sediments suggest the development of a small glacier in the Mulhacén cirque, which is reflected in several moraine ridges distributed across the cirque floor (Oliva and Gómez-Ortiz, 2012; Oliva et al., 2015).
LIA	From 1440 to 1710 CE a glacier existed in the Mulhacén cirque, and another glacier developed in the Corral de Veleta cirque until the mid-20 th century.	In the highest northern cirques between Mulhacén and Veleta peaks, historical sources and geomorphic evidence (i.e. moraines) shows evidence of the existence of several small glaciers that gradually disappeared during the 19 th and 20 th centuries (Gómez-Ortiz et al., 2001, Gómez-Ortiz et al., 2009, Gómez-Ortiz et al., 2018; Oliva and Gómez-Ortiz, 2012; Oliva et al., 2018a).

938

939 Table 2. Field data and characteristics of boulder samples dated with ^{10}Be in Sierra Nevada (*boulders are covered by snow thickness of 30 to
940 130 cm during 8 months per year). See text for details.

Sample	Latitude (°N)	Longitude (°W)	Elevation (m a.s.l.)	Thickness (cm)	Topographic Shielding factor	Snow Shielding factor
NAUT-1	37.0196111	3.33047222	2164	1.5	0.97832	
NAUT-2	37.0201111	3.33036111	2179	1.5	0.98414	
NAUT-3	37.02125	3.32977778	2211	1.5	0.98221	
NAUT-4	37.0225556	3.32866667	2263	3.0	0.97817	
SN-11-1	37.0595833	3.36569444	3076	2.5	0.80979	0.70067
SN-11-2	37.0605833	3.36747222	3061	2.0	0.80979	0.70067
SN-11-3	37.0606944	3.36897222	3095	2.0	0.80979	0.70067

941

942

943

944 Table 3. Analytical data and ^{10}Be sample exposure ages.

945

Sample	Quartz (g)	^9Be carrier solution (mg) ^a	$^{10}\text{Be}/^9\text{Be}$ ($\times 10^{13}$) ^b	[^{10}Be] ($\times 10^3$ at.g ⁻¹)	10Be age (CREp ^c) (ka)	10Be age (C-E ^d) (ka)	10Be age (CREp ^c) (ka) With snow shielding factor	10Be age (C-E ^d) (ka) With snow shielding factor
NAUT-1	18.32	98.9	23.3 ± 1.0	2530 ± 110	134.8 ± 10 (6.2)	130.4 ± 10 (6.1)		
NAUT-2	21.98	100.2	11.2 ± 0.58	1030 ± 53	53.7 ± 4.3 (2.9)	54.5 ± 4.3 (2.9)		
NAUT-3	22.53	100.5	9.62 ± 0.55	868 ± 50	43.8 ± 3.2 (2.3)	43.0 ± 3.6 (2.5)		
NAUT-4	22.62	100.7	28.6 ± 1.2	2580 ± 110	129.2 ± 8.9 (5.4)	125.4 ± 9.4 (5.5)		
SN-11-1	30.91	100.9	0.127 ± 0.052	8.4 ± 3.5	0.31 ± 0.10 (0.10)	0.28 ± 0.11 (0.11)	0.34 ± 0.12 (0.12)	0.32 ± 0.13 (0.13)
SN-11-2	32.79	100.7	0.285 ± 0.086	17.7 ± 5.3	0.60 ± 0.22 (0.21)	0.59 ± 0.20 (0.18)	0.72 ± 0.28 (0.27)	0.71 ± 0.22 (0.22)
SN-11-3	31.39	101.0	0.159 ± 0.050	10.3 ± 3.3	0.36 ± 0.09 (0.09)	0.35 ± 0.11 (0.10)	0.40 ± 0.11 (0.11)	0.38 ± 0.12 (0.12)
Blank		99.4	0.0450 ± 0.013					

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^a Carrier has a concentration of 3025 $\mu\text{g } ^9\text{Be/g}$.

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^b Sample $^{10}\text{Be}/^9\text{Be}$ ratios corrected for batch-specific analytical blank ratio of $(4.5 \pm 1.3) \times 10^{-15}$.

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^c ^{10}Be ages calculated with the online CREp exposure age calculator (Martin et al., 2017). 1sigma errors include analytical and production rate uncertainties; analytical errors shown in brackets.

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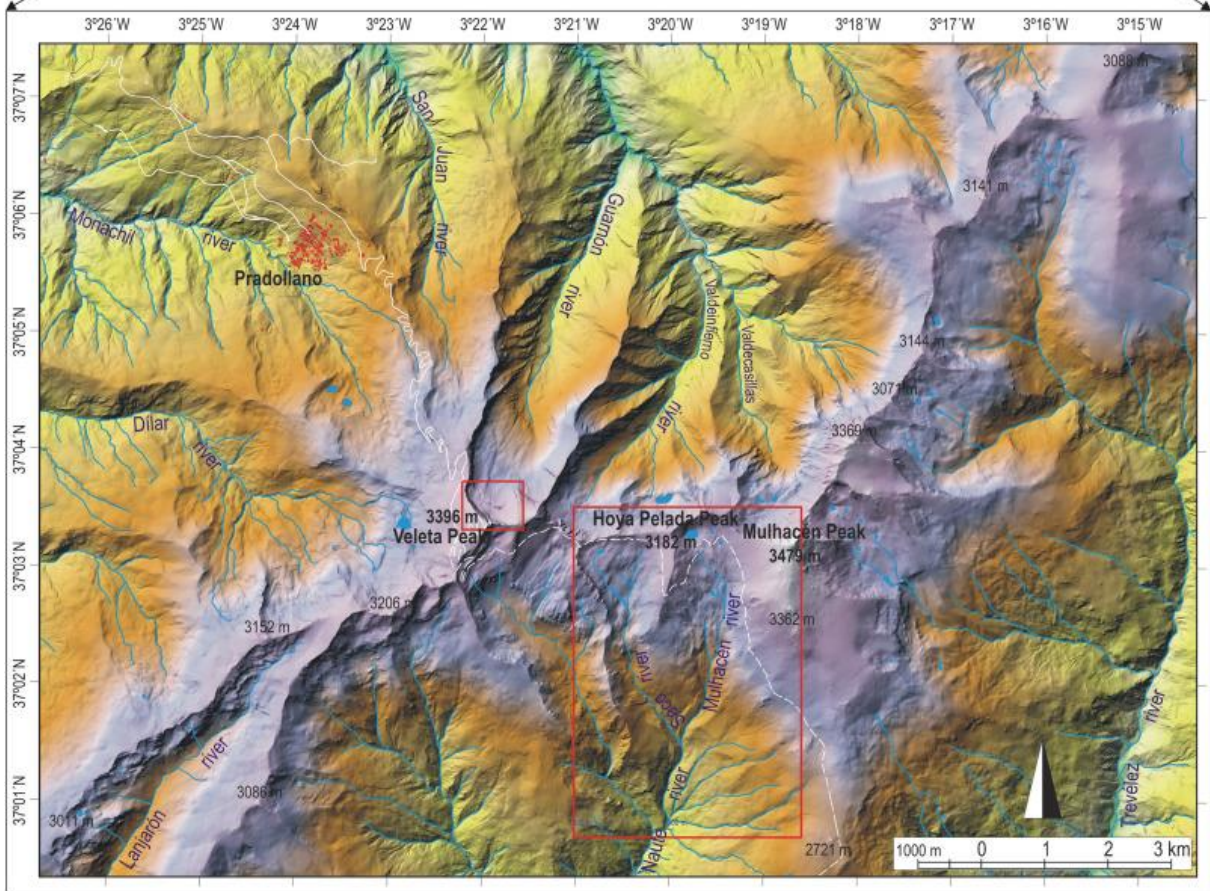
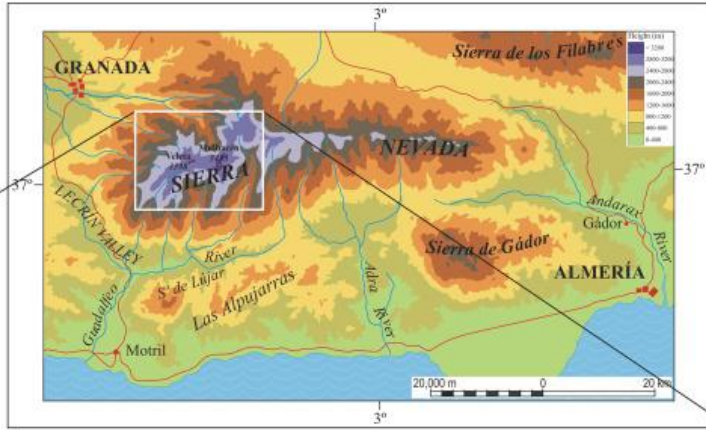
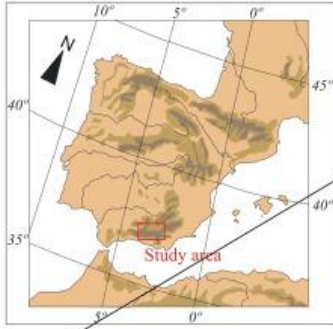
^d ^{10}Be ages calculated with the online exposure age calculator formerly known as the CRONUS-Earth online exposure age calculator version 3 (Balco et al., 2008). 1sigma errors include analytical and production rate uncertainties; analytical errors shown in brackets.

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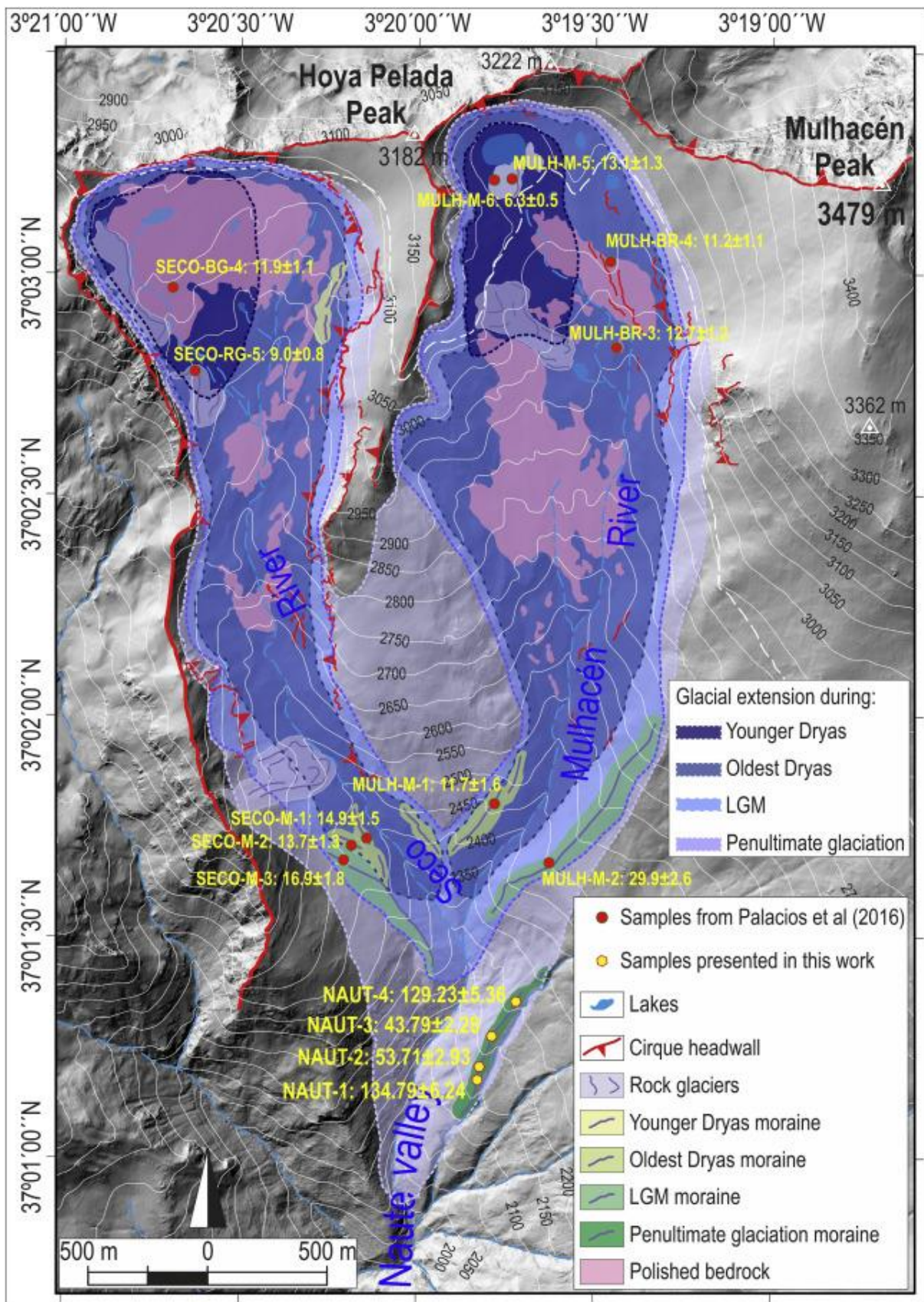
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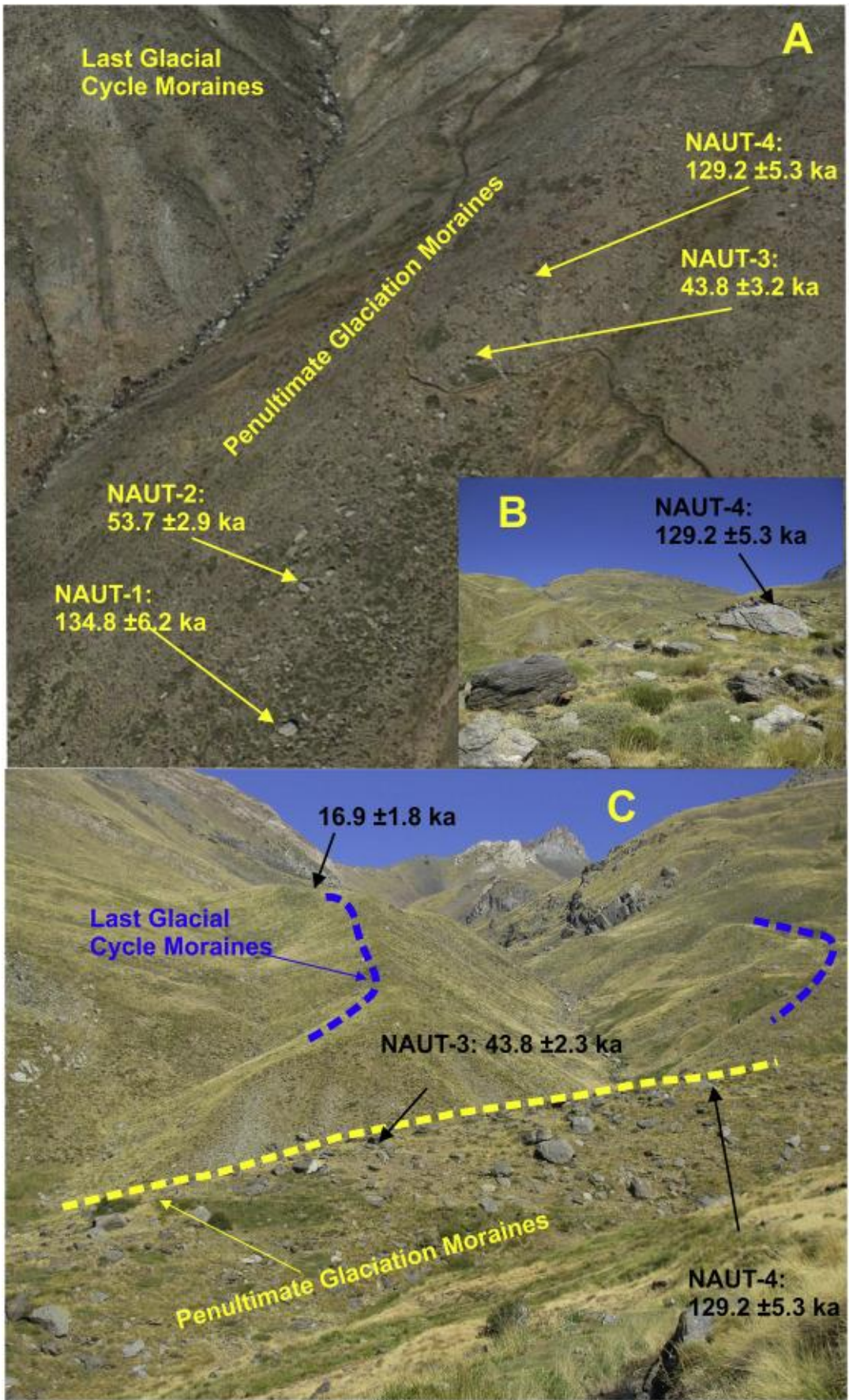
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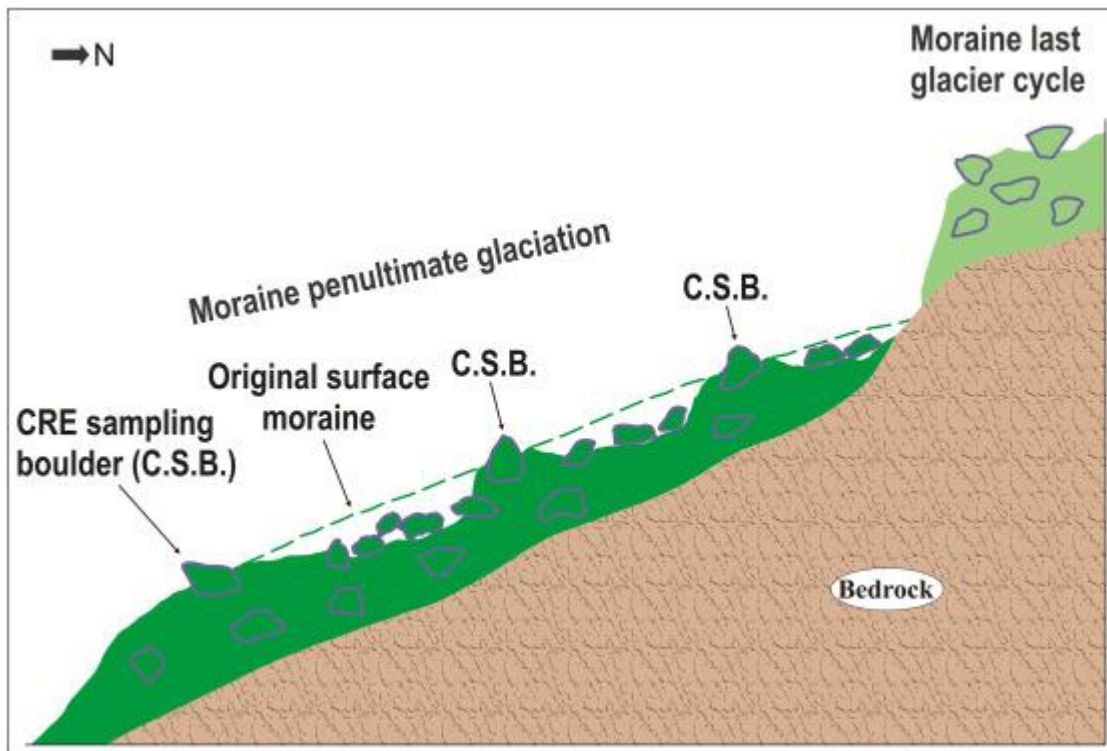
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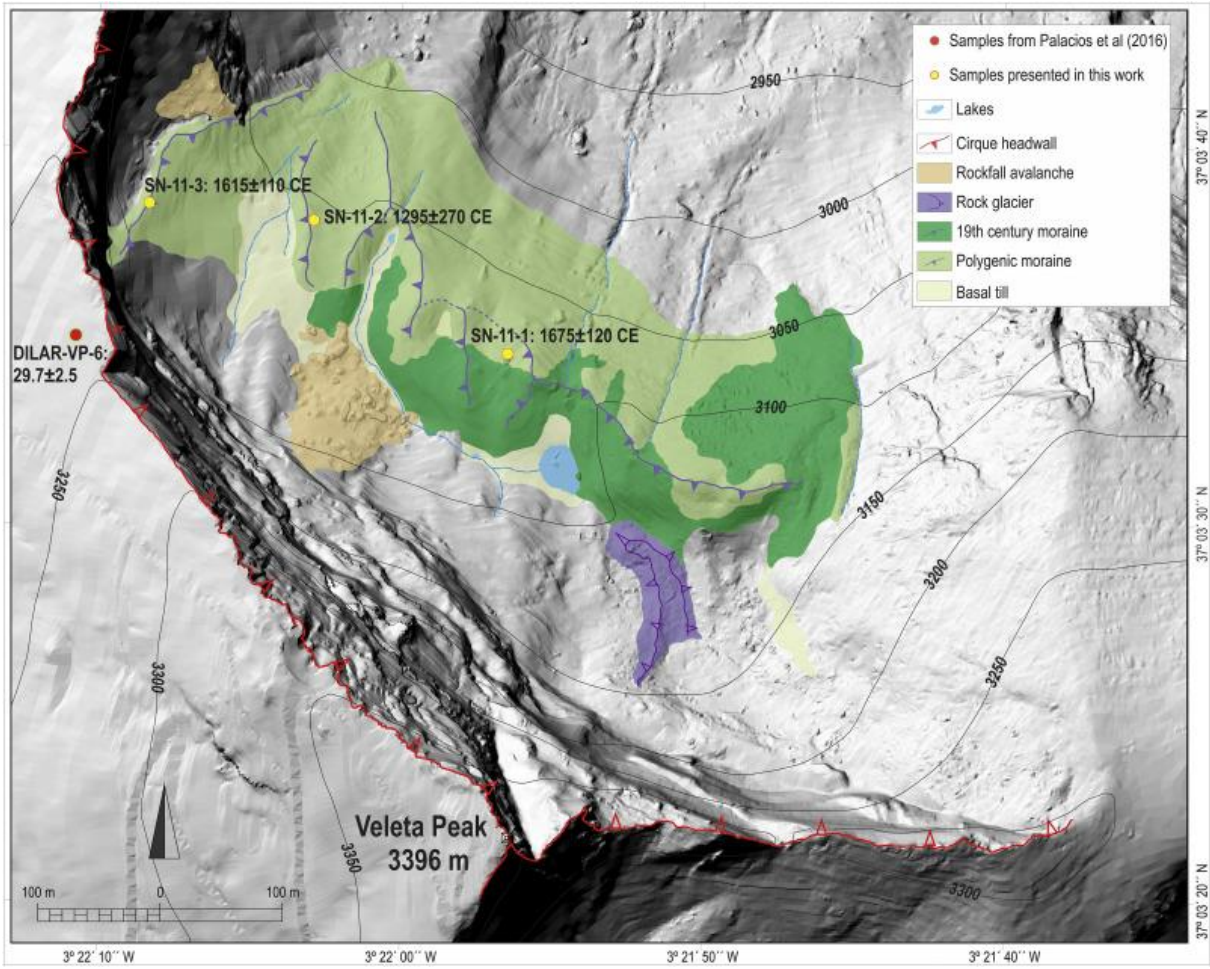
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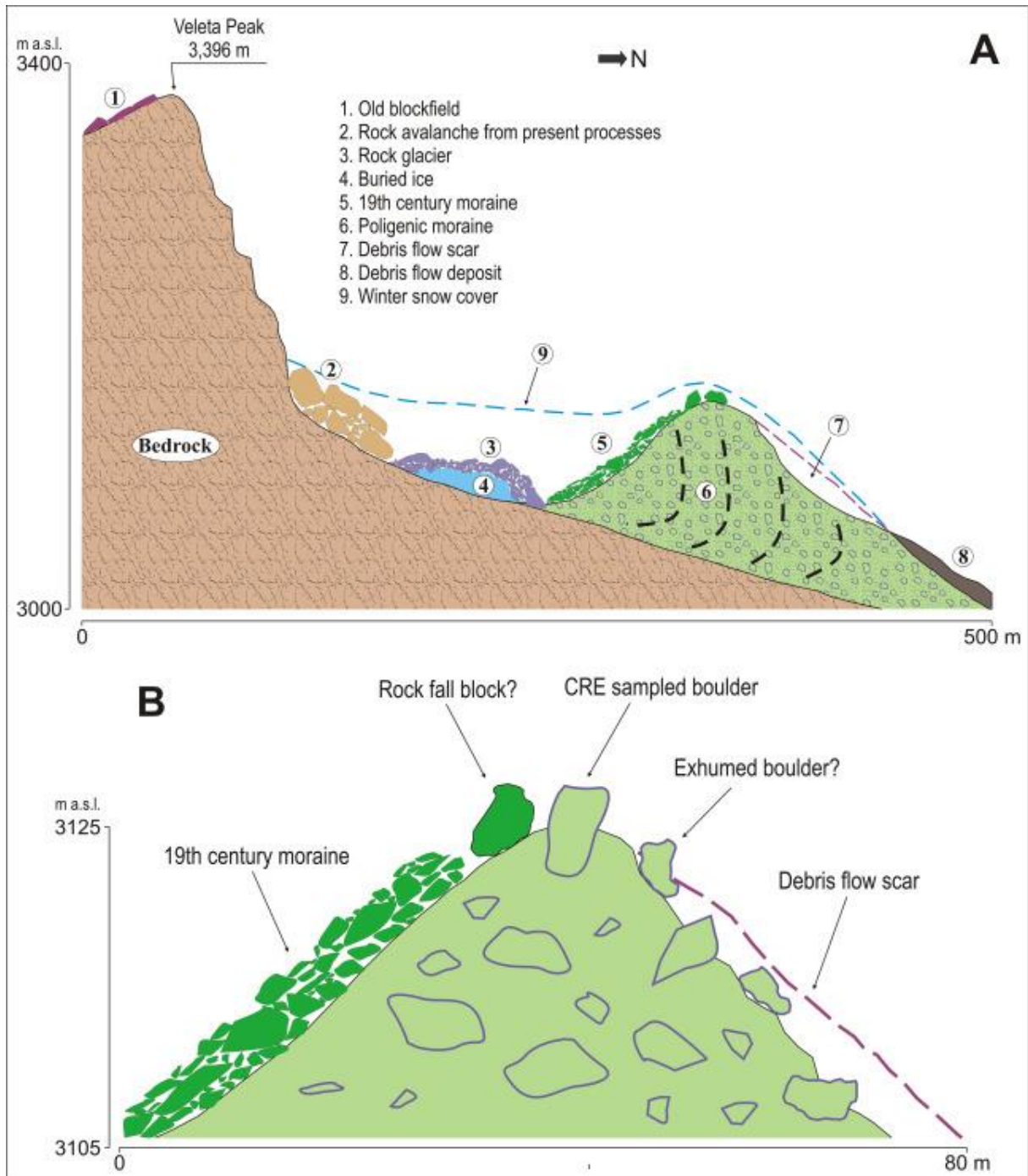
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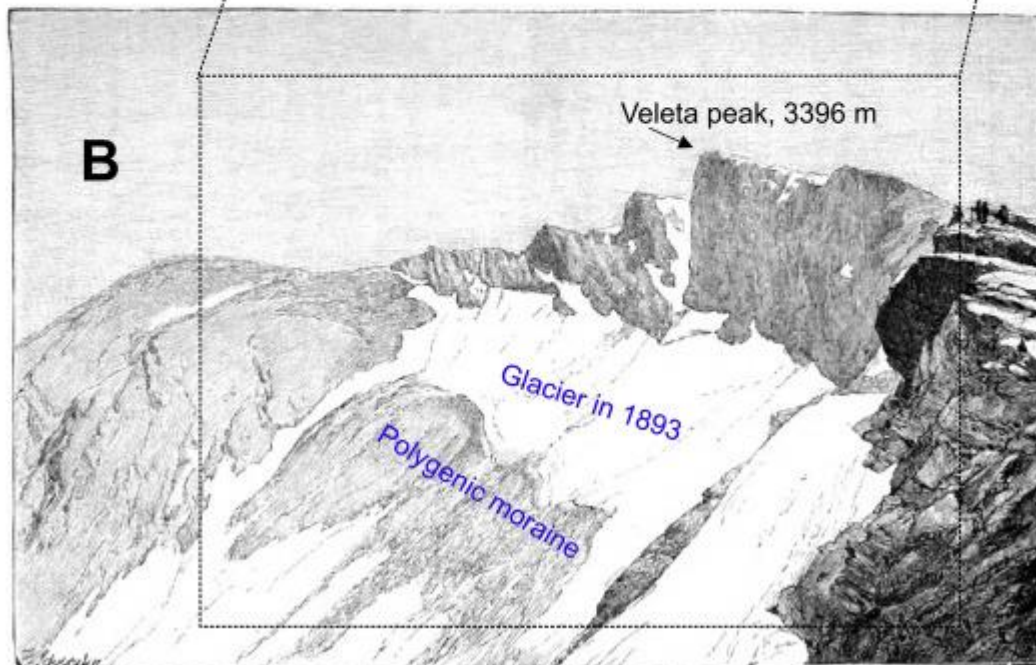
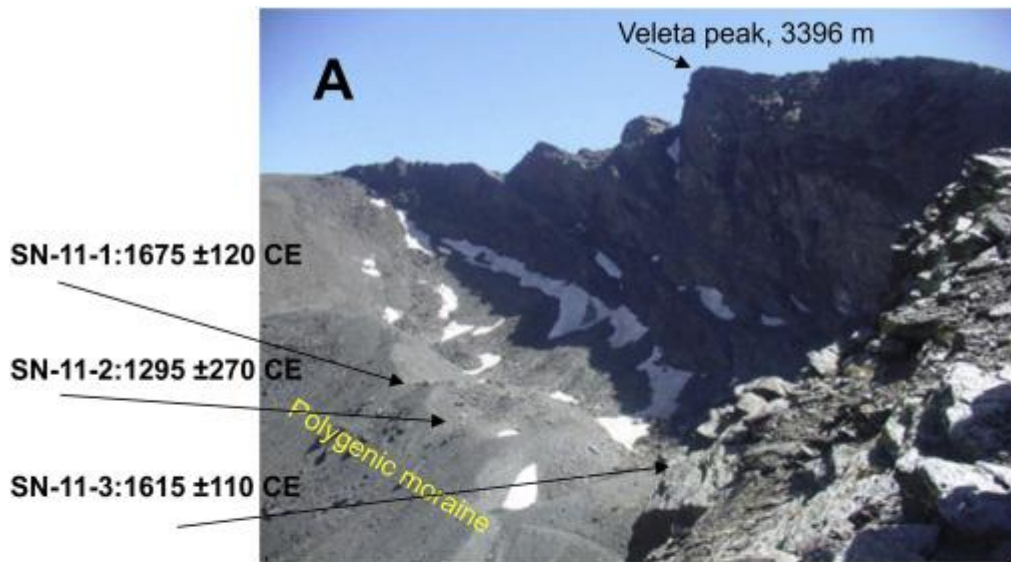
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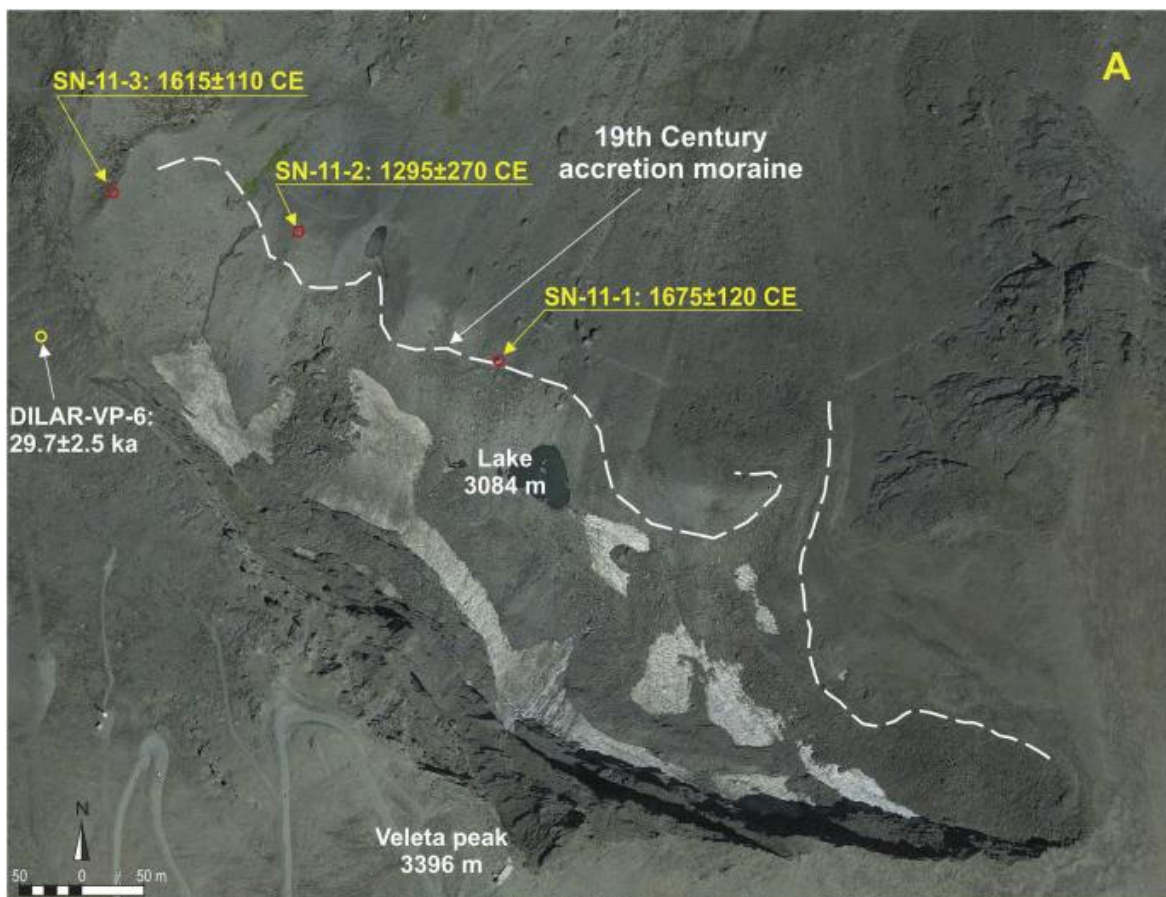
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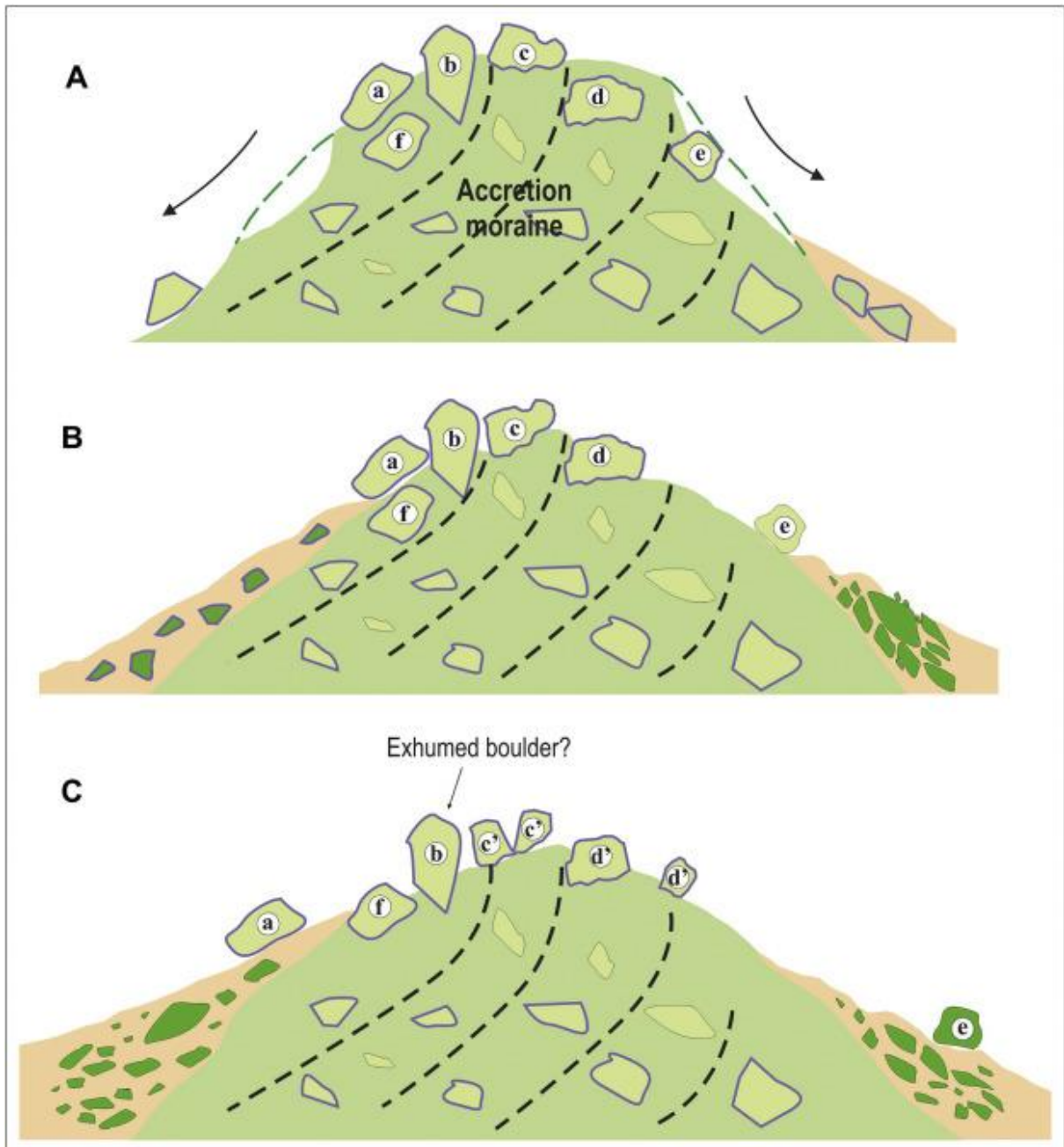
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972 **Figure captions**

973

974 Figure 1. Location of the case of study area, Sierra Nevada, in the south of the Iberian
975 Peninsula.

976 Figure 2. Geomorphological map of the Naute valley and location of samples. We show CRE
977 ages from this work (NAUT samples) together with those from [Palacios et al. \(2016\)](#).

978 Figure 3. Results of the Naute moraine samples of this work, together with pictures of some
979 of the sampled boulders and CRE data (including data from [Palacios et al. \(2016\)](#)). (A)
980 Penultimate glaciation moraine (source: Google Earth imagery), (B) Picture of the sampled
981 NAUT-4 boulder, (C) Panoramic view of the penultimate glaciation moraine next to the last
982 glacial cycle moraine.

983 Figure 4. Geomorphological sketch of the two moraines existing in the lower part of the
984 Naute valley, together with the location of the CRE Sampling Boulder (CSB).

985 Figure 5. Geomorphological map of the Veleta cirque and sample location. We show CRE
986 ages from this work (NAUT samples) together with those from [Palacios et al. \(2016\)](#).

987 Figure 6. (A) Schematic transect of the Veleta cirque including the different landforms
988 existing across the cirque floor, and (B) Sketch showing the geomorphological approach used
989 in the selection of samples in the Veleta cirque.

990 Figure 7. (A) Photo of Corral of the Veleta cirque from the West in September, 2016. (B)
991 Drawing according to the vision of [Bide \(1893\)](#), where the glacier seems still to cover its
992 maximum extension in 19th century.

993 Figure 8. (A) Veleta cirque with its 19th century moraine (source: Google Earth), (B) Vertical
994 view of the Veleta cirque floor from the Veleta peak, (C) Sampled boulder for SN-11-2, and
995 (D) Sampled boulder for SN-11-3.

996 Figure 9. Sketch representing the erosion processes affecting a moraine: (A) Intense mass
997 movements and erosion reshaping the main ridge during the paraglacial stage, (B)
998 Stabilization phase of a moraine at the end of paraglacial phase, and (C) Mature stage of an
999 old moraine after a long period of erosion.